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Assessing ionospheric influence on L-band SAR data: Implications on co-seismic displacement measurements of the 2008 Sichuan Earthquake.

D. Raucoules and M. de Michele

1
2 **Abstract**— Ionospheric contributions to the phase of L-band
3 **Synthetic Aperture Radar (SAR) signals put severe**
4 **limitations on ground displacement measurements retrieved**
5 **by either differential SAR interferometry (DinSAR) or radar**
6 **amplitude image offsets. Such contributions result in an**
7 **ionospheric phase screen (IPS) on the differential**
8 **interferogram and in directional fluctuations in the relative**
9 **position of azimuth pixels on offsets maps. In this article, we**
10 **propose a procedure for estimating and removing**
11 **ionospheric contributions to surface displacement**
12 **measurements derived from L-band SAR data. We test the**
13 **procedure on SAR data from the 28 May 2008 Sichuan**
14 **Earthquake.**

15 **The applied corrections allow both a clearer interpretation of**
16 **the surface rupture and a more accurate measurement of the**
17 **surface displacement, which has important implications in**
18 **earthquake modelling based on L-band SAR data.**

19
20 **Index Terms**—radar, interferometry, ionosphere,
21 earthquake

22 I. INTRODUCTION

23 **W**ithin the InSAR technique both the phase and the
24 amplitude of the backscattered radar signals can be used
25 for measuring earth surface displacements and
26 deformations. While DinSAR is based on the signal phase
27 difference between two radar acquisitions ([1],[2]) and provides
28 surface displacement values in the Line-of-Sight direction of the
29 satellite (LOS), the sub-pixel correlation technique measures the
30 sub-pixel offsets between two radar amplitude images both in the
31 azimuth and LOS directions of the satellite (e.g. [3],[4]). The
32 former technique is as accurate as a fraction of the employed
33 radar wavelength and is sensitive to mm to dm surface
34 displacement. The latter

35
36 technique is generally sensitive to ground displacements larger
37 than 0.1 pixels, which is about 50 cm in the azimuth direction for
38 a space-borne radar sensor such as the Phase Array L-band
39 Synthetic Aperture Radar (PALSAR). These two techniques are
40 complementary, particularly when LOS deformation gradients
41 larger than one quarter of the wavelength per pixel cause
42 interferometric signals to de-correlate. This might occur close to

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43 a seismic rupture, such as the ~270 km long Sichuan earthquake
44 rupture where co-seismic slip reached up to 8 metres [5].

45 While L-band SAR signals are of particular interest in studying
46 earthquakes as it is less affected by temporal canopy changes
47 than C-band, it could be severely

48 affected by ionospheric heterogeneities occurring during the
49 synthetic aperture calculation (e.g. [6],[7]). Due to the
50 “dispersive” nature of the medium, the ionosphere refractive
51 index depends on the inverse of the square of the electromagnetic
52 frequency employed. Therefore, L-band SAR data are more
53 affected than C-band SAR.

54 The ionosphere influence on the SAR signal affects both azimuth
55 sub-pixel offsets and differential interferograms. The first bias
56 results from directional fluctuations in the relative sub-pixel
57 position of azimuth pixels, already reported in literature as
58 azimuthal “streaking” ([7]). The second bias results from relative
59 lengthening of the wave paths between two radar acquisitions
60 affecting the interferometric signal. As reported by recent studies
61 based on L-band InSAR on the Sichuan earthquake ([8],[9]),
62 ionospheric influences on the SAR signal appear to introduce
63 several difficulties for the retrieval of surface deformation from
64 both sub-pixel offset of radar amplitude images and differential
65 interferometry.

66 In this paper we focus on the ionospheric influences on the
67 SAR signal, assess their impact in the presence of co-
68 seismic surface displacement and try to propose a method
69 to estimate and remove their contributions both to sub-
70 pixel offset and to interferometric phase. We apply the
71 method to the 2008 Sichuan earthquake surface
72 displacement measurements.

73 II. IONOSPHERIC EFFECTS ON SAR DATA

74 A. Interferometric phase

75 The impact of the ionosphere on the interferogram is caused by
76 the relative variation of the refractive index of the medium
77 between the two radar acquisitions. The resulting propagation
78 lengthening produces an interferometric phase shift.

79 This phase shift is related to the electron density variation, n_e at
80 height h in eq.1. For a nadir-looking radar [11]:

$$\Delta\Phi \approx -\frac{4\pi}{c_0} \frac{40.28}{f} \Delta TEC \quad (1)$$

$$\text{with } TEC = \int_0^H n_e(h) dh$$

86 where ΔTEC is the variation of the Total Electron Content
87 (TEC), c_0 the speed of light, f the signal frequency (Hz).

B. Azimuth streaking

According to [7] atmospheric impact on C-band and L-band InSAR results from radar signal phase modulation due to spatial variation of the ionospheric propagation conditions during the aperture time.

[11] derived the relation between ionospheric contribution to the azimuth offset (Δx_{iono}) and the ionospheric contribution to the interferometric phase, which can be expressed as:

$$\Delta x_{iono} = \alpha \frac{\partial}{\partial x} (\Delta \Phi_{iono}) \quad (2)$$

Based on this relation we can estimate the interferometric phase correction, or IPS, starting from the azimuth offsets values [7]. We can then remove the IPS from the interferogram to enhance the coseismic deformation measure.

III. AZIMUTH CORRECTION

We observe that the azimuth streaks on the sub-pixel offset map show up with a preferential direction (figure 2a). Azimuth offsets are estimated on co-registered images (co-registration based on the adjustment of a bilinear model) of the interferometric pairs and therefore with identical geometry. In such conditions possible residual topographic effects are very limited (which is not the case with slant range offsets for with stereoscopic effect is not negligible even with perpendicular baselines of some tens of metres).

Over the Sichuan earthquake area the azimuth streaks direction is constant over a large spatial and temporal scale, at least during the concerned period (e.g. [8],[9]). The direction of the streaks seems to be constant for a given geographical area across different radar tracks. However, [7] who worked on polar areas noticed, in certain cases, along-track variation of this direction. The influence of the position respect to the magnetic poles has to be investigated for other test sites.

On the Sichuan area, the streaks strike \sim N115E, while the earthquake ruptures strike \sim N40E [5]. According to [14], we can observe that South China is located in an area affected by a strip of high electron density (related to the location of the geomagnetic equator) which main orientation roughly corresponds to \sim N115E. That could explain the direction of the streaks and the high values of ionospheric effects on the area.

Based on this peculiarity we improve the methodology firstly proposed by [7] by taking into account the spatial evolution of the azimuth streaks amplitude along their length over the entire radar image width.

In order to reduce the azimuth streaks, [8] proposed to cut the azimuth offset map into three sub-images within which the azimuth offset correction is approximated by a constant value along the streaks direction (*i.e.* the correction is constant by segments corresponding to the subimages). This method provides a satisfactory correction to highlight the surface trace of the earthquake rupture and does not affect coseismic offset values in the near field. However, this approximation yields residual discontinuities at the sub-images boundaries.

Among the 1D low pass filters that could be used for this purpose, we propose to use single polynomial fits. In this paper, we approximate the azimuth streaks amplitudes by a third degree polynomial along the streak direction. After rotation of the image in order to align horizontally the streaks, each line is replaced by its third degree fit. This approximation fits the trend well enough

to remove most of the azimuth streaks without affecting high-frequency small-scale offsets, such as near field offsets due to the earthquake rupture. We test the methodology on two different ALOS PALSAR tracks (table 1) acquired over the Sichuan area. We first test the methodology on a radar track less affected by coseismic deformation (figure 2a) then we apply it to enhance the coseismic rupture on a different track (figure 3a/b). We assume that the computation of the α coefficient on the track less affected by deformation is more reliable as the offset and phases are mainly related to ionosphere and not deformation. Note that α only depends on sensor parameters, so α is the same for both frames.

In both cases, we compute sub-pixel offsets maps on full resolution amplitude images by using the GAMMA routines (<http://www.gamma-rs.ch/>), from which we subtract a linear offset ramp (figure 2a). The linear offsets ramp is due to image co-registration procedure and residual uncompensated orbits. The ramp does not mask deformation but can be considered as a bias. On the other hand, the objective of removing a ramp from offsets is to obtain a result comparable to InSAR. In fact, InSAR is also biased by a phase ramp on scales larger than the image coverage. Such effects systematically affect InSAR results [13].

Figure 3a) and 3b) show an example of correction applied to the azimuth offset map on track 473, concerned by the seismic slip with values of up to 5m. We can notice that the deformation was initially masked by the ionospheric contributions to azimuth offset (figure 3a). After correction, the coseismic rupture is enhanced and it can be mapped. Also we can retrieve the azimuth component of the near field coseismic offset (about 1 pixel in the azimuth direction, *i.e.* 3.6 metres), which is crucial for inverse modelling of the earthquake

IV. COMPUTATION OF THE PHASE DERIVATIVE

As pointed out in section 1, the contribution of the ionosphere to the azimuth offset can be associated to the along-track derivative of the interferometric phase. We will use this information to calculate the IPS and remove it from the coseismic interferogram.

As a prior processing step to estimate the phase correction, we compute the phase derivative on an extended area of track 471 where we infer no major surface deformation has occurred. In order to reduce noise, we applied a 20 pixel mean filter in the columns direction. Then, we compute the derivative by using eq.3. With this formulation, the derivative respect to the line index i for a given pixel (i,j) can be estimated on the complex interferogram without unwrapping.

$$\frac{\partial}{\partial i} (\Phi_{i,j}) = W(W(\Phi_{i+1,j}) - W(\Phi_{i,j})) \quad \text{with } W(x) = x [2\pi] \text{ and } W(x) \in [-\pi \text{ rad}, +\pi \text{ rad}] \quad (3)$$

The benefits of the phase derivative image are twofold. Firstly, it allows us to confirm the validity of the azimuth offset correction estimation. In fact, similarities between the pattern of azimuth offset correction estimation (figure 2b) and the pattern of the phase derivative (figure 2d) confirms the validity of the former, as stated in eq. 2. On the other hand, the comparison of both results, allows us to estimate the proportionality coefficient α (eq. 2) necessary for IPS estimation. We examine both the standard deviations (table 1) and the scatter plot on a selected area of track 471 (figure 2d).

Considering the linear relation between the two datasets, the α value can be estimated as:

$$\alpha = \frac{\text{std}(\Delta x_i)}{\text{std}\left(\frac{\partial}{\partial i}\phi\right)} = 30.8 \text{ pixels/radian} \quad (4)$$

$$\beta = 1/\alpha = 0.032 \text{ rad/pixel}$$

These parameters depend on geometric characteristics of a given sensor in a given mode [11]. Therefore, once α and β are estimated on a given radar track/frame, one can use them for correcting other radar frames acquired by the same sensor. This point is of particular importance as we should not compare the interferometric phase derivative with the azimuth offset correction over an area that is affected by a high surface deformation gradient since, in this case, the phase derivative would be affected by surface deformation and the estimation of α would be biased.

V. INTEGRATION OF THE AZIMUTH CORRECTION: THE IPS

Now, we calculate the IPS and we correct the differential interferogram on track 473. The IPS is the results of along-orbit integration of the ionospheric contribution to the azimuth offsets (obtained in section III), converted into the phase screen. The conversion from azimuth offsets to the phase screen is obtained by dividing offset values by coefficient α .

$$I_{i,j} = \sum_i \Delta x_{i,j} \quad (5)$$

$$\Delta\Phi_{iono i,j} = \frac{1}{\alpha} I_{i,j}$$

Where the I are the results of the integration, x is the ionospheric contribution to azimuth offsets, i and j are line and column indices respectively.

Figure 4a shows the extracted ionospheric contribution to the interferometric phase for track 473. In this case study, we can notice that the total ionospheric contribution to the interferometric phase corresponds to ~ 4.0 radians, which makes ~ 0.6 interferometric fringes or ~ 7.5 cm apparent surface displacement in the LOS direction for PALSAR. For track 471 (figure 2e and 2f) the IPS undulation is up to ~ 18 radians (i.e. ~ 3 fringes) equivalent to 34 cm displacement in LOS. Such a phase contribution can severely affect physical interpretations of the earthquake surface deformation based on L-Band interferograms over the mid-to-large scale deformation field, more precisely for wavelengths equal or larger than about 25 km^{-1} . A more detailed study should be carried out to investigate the consequences of the ionosphere on shorter wavelengths.

Figures 2b-2f and 4b-4c compares the differential interferograms before and after correction. We can notice that the total deformation pattern is different. For track 471 we can observe a clear improvement of the interferogram such as decreasing of the fringe number and regularisation of the fringe pattern. A large wavelength bias of about 1 fringe still affects the corrected interferogram. It is probably due to uncompensated orbital ramp or a residual tropospheric contribution. At this stage, a quantitative validation on the improvement made on the coseismic interferogram is difficult to carry out. A validation would require a dense ground measurement network covering the entire study area. At the moment we are writing this manuscript, the GPS coverage on the area is not enough dense to allow a consistent validation ([15]).

VI. DISCUSSION

In this case study, the highest co-seismic slip (up to 8 metres) is located within ~ 15 km of the rupture. Thus we assume that the third-order polynomial used to calculate the ionospheric contribution to the azimuth offsets has a minor influence on the near-field co-seismic displacement values measured by offsets (i.e. close to the rupture).

On the other hand, far-field deformations are usually smaller and of longer wavelengths. In this case, the ionospheric contribution to the interferometric phase should not be neglected for a correct interpretation of the surface deformation.

Moreover, we have to notice that in another case where deformation field produces long wavelength offsets in the same direction as the azimuth streaks, our methodology might result in an underestimation of the surface displacement as deformation signals would be more difficult to separate from the ionospheric contribution. On the other hand, given the nature of the ionospheric influence on the SAR signal, i.e. it concerns the derivative of the interferometric phase, independent TEC measures (such as by GPS, for instance) might not be adequately dense to resolve the mid wavelength ionospheric derivative and thus they would not be helpful in modelling and removing the ionospheric contribution to the azimuth offsets.

VII. CONCLUSION

In this paper, we proposed a procedure for extracting ionospheric contributions to the SAR signal and we apply it to improve earthquake measurements based on PALSAR L-band SAR data over the Sichuan earthquake area. We used both sub-pixel correlation of radar amplitude images and radar interferometry. Based on the directionality of the azimuth streaking we defined an adaptive directional filtering method and approximated the ionospheric contribution to the azimuth offset. We then used this estimate to assess and remove the ionospheric contribution to the interferometric phase. The two following observations resulted from the presented study. Firstly, although they were initially severely affected, the azimuth sub-pixel offsets can be used both to precisely map the earthquake rupture and to measure the coseismic displacement in the near field. Secondly, the ionospheric contribution to the interferometric phase (i.e. the IPS) can reach up to ~ 15 radians, equivalent to ~ 28 cm apparent LOS displacement for ALOS PALSAR. This has important implications when using L-band interferometry to model the earthquake cycle.

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TABLE I
ALOS PALSAR ACQUISITIONS USED FOR THIS STUDY

Track	Date	Mode	Pixel size Range/Azimuth (m)
471	29/02/2008 31/05/2008	Ascending	4.68 / 3.15
473	17/02/2008 19/05/2008	Ascending	4.68 / 3.15

TABLE II
STANDARD DEVIATION OF THE PHASE DERIVATIVE AND THE AZIMUTHAL CORRECTION ON THE SOUTHERN
AREA OF TRACK 473

	standard deviation	mean
InSAR Phase derivative	0.013 rad	0.0024 rad
Azimuth correction offset	0.37 pixels	0.06 pixels

Figure 1: Area of interest. Locations of PALSAR acquisitions used are identified by the dotted rectangles. WF= Wenchuan Fault; BF= Beichuan Fault; PF= Guanxian-Pengguan Fault (modified from [10]).

Figure 2 a) azimuth sub-pixel offset map (track 471). Azimuth streaks are clearly visible. b) Ionospheric contribution to the azimuth offsets after directional polynomial fitting and linear trend removal c) Interferogram (track 471). Patterns with several fringes orientated in the 'streaks' direction are visible d) Along-orbit phase derivative (track 471). We observe the similarity with the azimuth streaks. The scatterogram between b) and d) data is plotted. It is consistent with a β coefficient of 0.032 rad/pixel (white line). e) IPS computed for track 471. f) corrected interferogram. The dashed area corresponds to a noisy area on the azimuth offset image and therefore irrelevant correction.

Figure 3 a) azimuth offset map for track 473. b) Azimuth offsets map corrected for the ionosphere. The white line is the Sichuan earthquake surface rupture measure on the field (modified from [5]).

Figure 4 a) Ionospheric phase screen for track 473. b) Differential interferogram for Track 473 c) Corrected differential interferogram for track 473.







